



A plan for the VLT

Grenoble - 15/16 January 2014



2 Timescales

- ◆ Preparing for GRAVITY and MATISSE
- ◆ VLT in the ELT Era (ESO VLT White book)

Requests

- ◆ Prepare the infrastructure for GRAVITY and MATISSE
- ◆ Improve VLTI performance:
 - Robustness to weather conditions: adaptive optics on ATs (NAOMI)
 - Robustness to vibrations in the UTs: active vibration control strongly recommended by STC
 - Sensitivity and spectral resolution: provide fringe tracking
- ◆ Improve the access to the VLTI community:
 - Science ready pipelines: ESO deliverables
 - Phase 3 (science ready) data product repository: ESO repository (JMMC repository: OIDB).
 - Production of phase 3 products by ESO: not resourced
 - Providing reconstructed images: community
 - Promoting the use of VLTI: community (ESO)



An important effort

NOW	Status
MIDI	Removed in P94
PIONIER	Removed in P94
AMBER	Stays until:
FINITO	Stays if AMBER stays
NAOMI	PDR end of 2014 ?
VIBRATION	No project yet
GRAVITY	PAE + shipment end of 2014
MATISSE	GRAVITY + 1yr ?
PRIMA	Decision February 2014
2GFT	Delayed 2016 ?

Need for a holistic technical implementation plan



Details

◆ PRIMA-Astrometry:

- Commissioning process interrupted 2011;
- Switched to “engineering mode” to enable “experimental” astrometry and evaluate performance;
- Conclusion of engineering period: system as designed can’t deliver the requested performances for science case. Need an improved baseline calibration.
- Gate review principle endorsed by ESO management and approved by STC: 29/30 january
 - Non-ESO board;
 - Validity of PRIMA recovery plan evaluated (technical, resources, managerial)
 - Pertinence of science case in the exoplanet context.
- Based on review, the ESO management will take the decision to continue or not PRIMA-Astrometry

Details

◆ PIONIER:

- Recommended by STC, principle of its extension approved by Paranal director pending status clarification and agreement with IPAG.
- Technical solution to relocated PIONIER proposed by Paranal: no resources.
- Phase 1: If implemented will remain one/two years with current status.
- Phase 2: Service mode ?
- Phase 3: Upgrade (J, medium resolution)?

Details

◆ AMBER

- Can stay after the lab modification;
- Aging instrument
- Drop in number of night requested (end of its life as it is ?)
- Will it be competitive at the time GRAVITY comes is offered?
- Spot could be used for J,H + 4T + spectral resolution ?
- What do we do with AMBER ?

Details

- ◆ NAOMI: PDR in view (end 2014) slow pace
- ◆ Vibrations control: Identified as priority by STC. Currently no formal project at ESO
- ◆ 2nd generation fringe tracking: Postponed to 2016. Serious problem for MATISSE.



The future of VLTI

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VLTI strengths

- ◆ The best angular resolution accessible in Europe (for a long time);
- ◆ After a huge investment it is finally working well
- ◆ AT sub array
- ◆ Two new instruments coming;
- ◆ ESO (service mode, archives, maintenance, stability)
- ◆ An untapped potential in the study of stellar environments and AGN;
- ◆ An untapped potential in the study of multiplicity (massive stars, brown dwarf, ysos)
- ◆ An untapped potential in the study of temporal variability (ysos, novae, evolved stars Bes)
- ◆ An untapped potential for creativity (GRAVITY)?

VLTI strengths

- ◆ Workhorses topic that appear in almost all prospective exercises e.g Astronet, NASA, US decadal (YSOs and AGNs) with growing interests from the community.
- ◆ A very active community, well organized (schools, EII ...)
- ◆ A unique potential for community growth (e.g PIONIER days)
- ◆ Entire bandpasses unexplored (can be opened for 1 Meuro);
- ◆ JMMC: a community center that develops useful software for the community with high reactivity.
- ◆ An untapped potential of synergy with ALMA
- ◆ Optical interferometry is a serious contender to become the post-ELT ground based optical facility (ok we are speaking about decades ...)

VLTI weaknesses

- ◆ The community is small and partly interferocentric;
- ◆ The accessibility to reduced data is a problem (prebiscit of PIONIER's black box + service package): 10 years after this is a problem.
- ◆ Long-baselines lacking: deficit of resolved stellar surfaces
- ◆ The funding context (ELT) is hostile
- ◆ ESO lack of flexibility
- ◆ The ESO current organisation context (no more Paranal VLTI, probably no more Garching VLTI) weakens the possibility to develop the VLTI in a holistic way (fragmentation)

Plans for the future

- ◆ Make GRAVITY and MATISSE a success
- ◆ Construct a real lobbying power (science based)
- ◆ Need for the presence of the VLTI in the VLT whitebook (“VLT in the ELT era” status not clear)
- ◆ Explore ways to accompany the VLTI user:
 - ◆ ARC-like center ? JMMC, EII funding
- ◆ High dynamic binary explorer (4T, L band (nulling?), precision closure phase)
- ◆ J, H with medium resolution imager (molecular bands in cool atmospheres, envelopes, Paschen Beta spatially resolved mass accretion/mass loss, radius-luminosity)
- ◆ Very high spectral resolution in the visible, 4T: link with asteroseismology?,
- ◆ 6 telescope imager? limitations in imaging power clearly visible, the necessity to tackle temporary varying processes.
- ◆ Heterodyne interferometry at Paranal ?

A POSSIBLE roadmap

- 2014: Start community discussion on VLTI “nodes”? (EII)
- 2015: Stellar physics 2.0 workshop (short wavelengths oriented) with VLTI whitebook in mind?
- 2016: removal of AMBER
- 2016: decision on fringe tracker solution
- 2018: decision on PIONIER++ (J,H Medium resolution)
- 2018: decision on Binary Hunter (BROWNIE)
- 2020s GRAVITY/MATISSE return of experience in imaging: should we go 6T?



Discussion

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Questions ?

- ◆ **Period 94 shutdown tbc**
- ◆ What should we do with AMBER?
- ◆ **How does the community structures itself (inspiration from ALMA?):**
 - to interact with ESO (e.g: future of operations, decision on AMBER)
 - to prepare accompany GRAVITY and MATISSE: ARC (?)
 - Xavier Haubois (JMMC)
 - What European initiatives ?
- ◆ How to you see the future of VLTI ?
 - ◆ A possible Roadmap:
 - 2014: Start community discussion on VLTI “nodes”? (EII)
 - 2015: Stellar physics 2.0 workshop (short wavelengths oriented) with VLTI whitebook in mind?
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 - ◆ Connexion with PFI ?